

# Social Network Analysis of the Planetary Data System

*Final Presentation*

Kate Martin  
Mark Avnet

May 9, 2006

ESD.342  
Advanced System Architecture

# Overview Planetary Data System

## NODES/SUBNODES/DATA NODES

Function

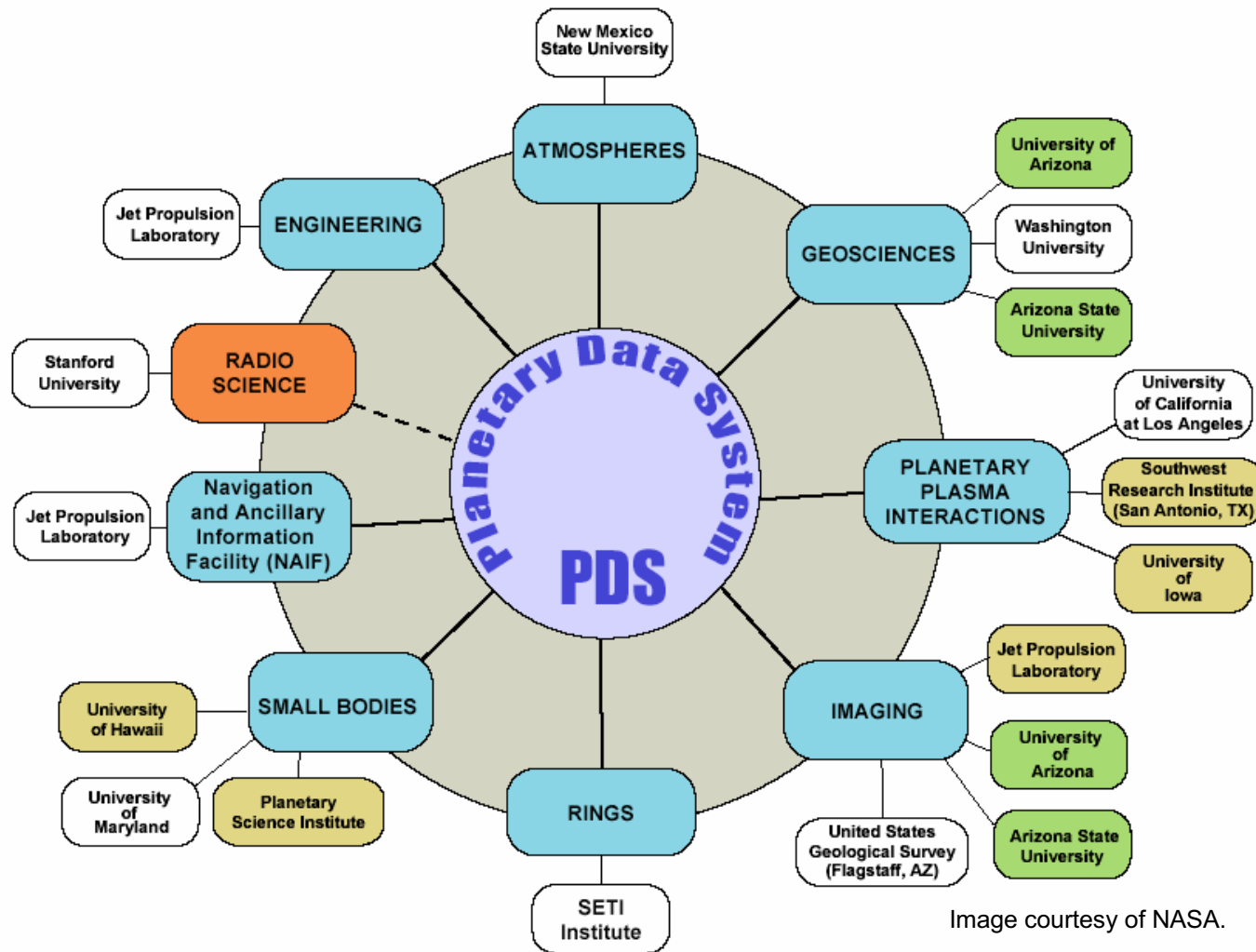
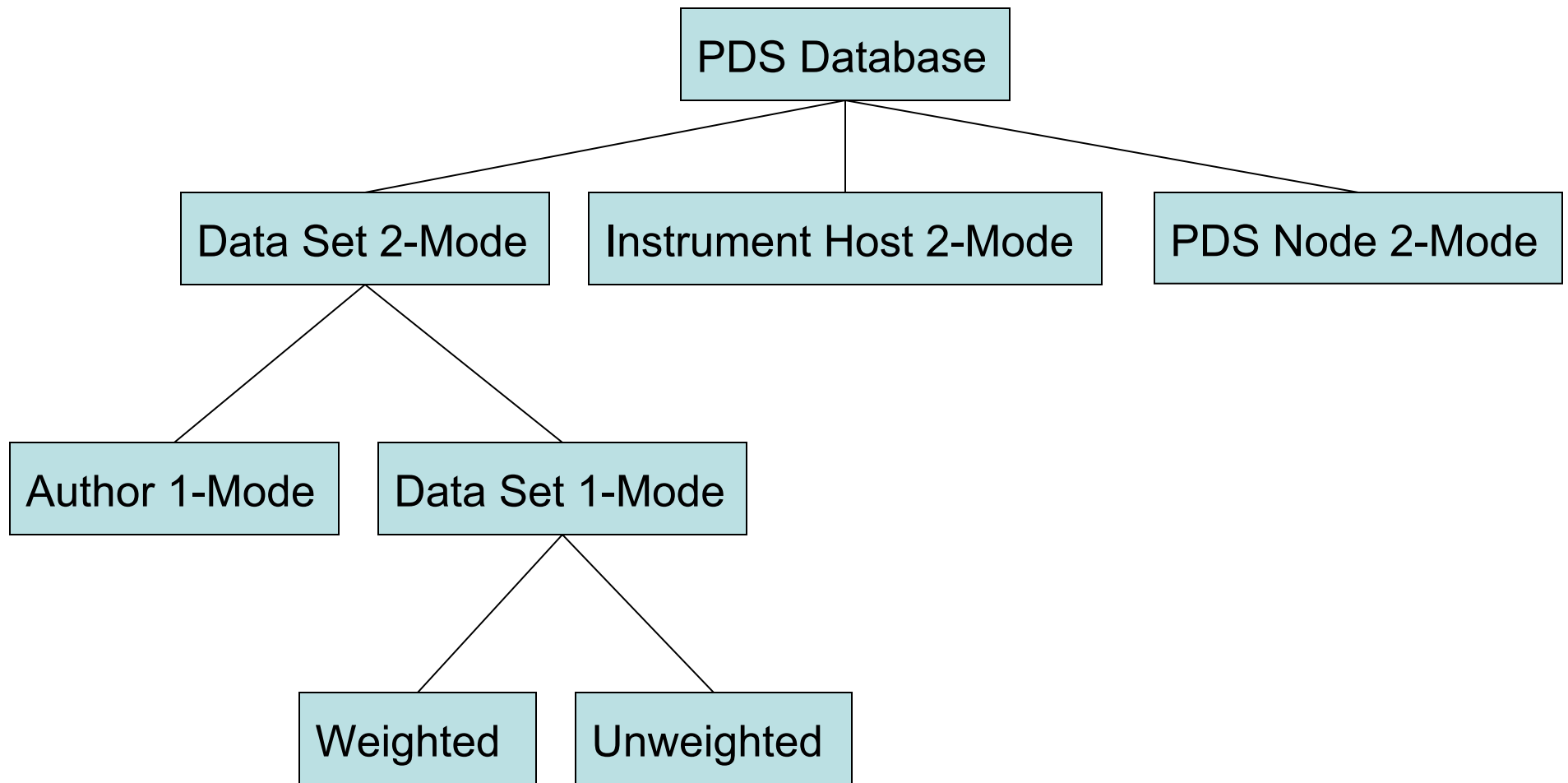


Image courtesy of NASA.

# Decomposing the PDS Database into 3 Bipartite Networks ...



... Yields 12 separate 1-Mode networks  
for comparative analysis

# Newman-Type Table for the 12 PDS Networks

Type of Node	Network	Weighted?	$n$	$m$	$\langle k \rangle$	$l$	$\log n / \log \langle k \rangle$	$\alpha$	$C^{(2)}$	$\langle k \rangle / n$	$r$
<b>Authors as Nodes</b>	PDS Nodes	No	439	27493	125.3	1.886	1.260	-0.31	0.981	0.285	0.80
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- **Network-Specific Results**

- Small-magnitude  $\alpha$  values, but the regions that follow a power law are small for many of the networks
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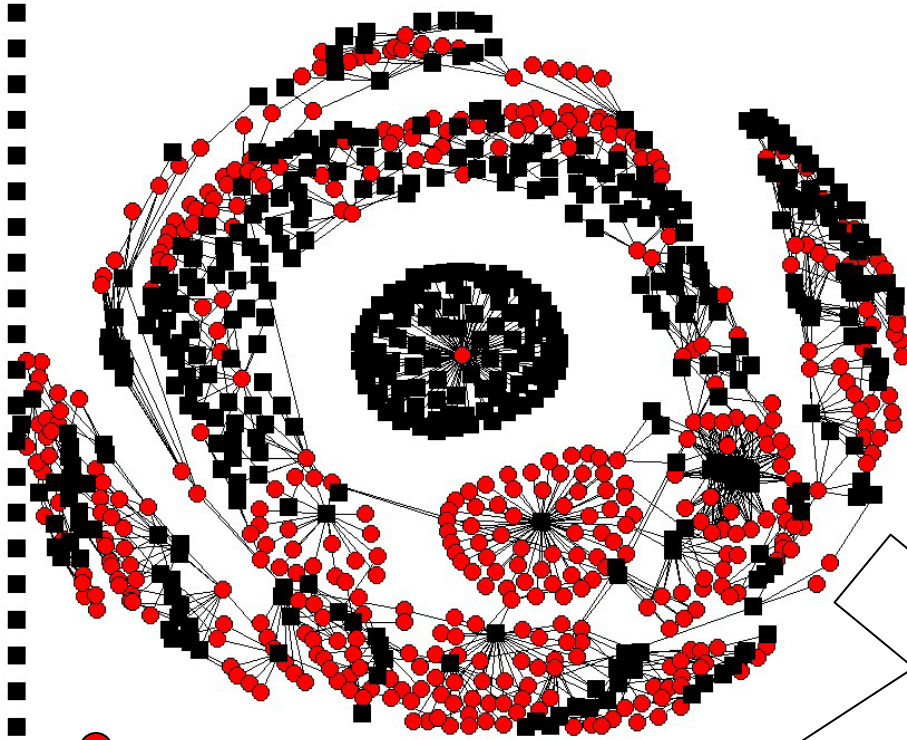
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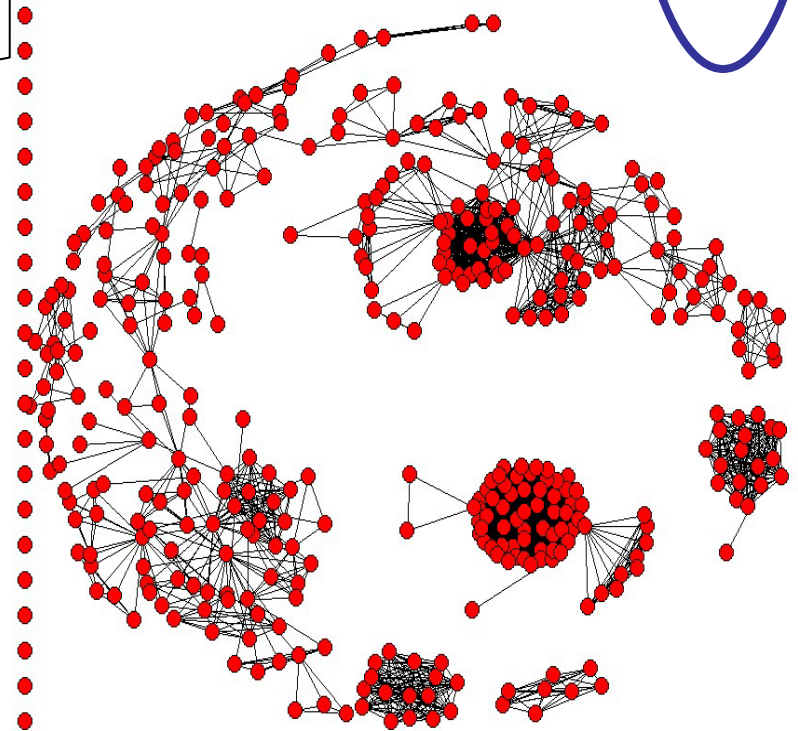


- Authors
- Data Sets

Affiliation networks lead to triangle motifs and high clustering coefficients

Network of authors with data sets as edges

Type of Node	Network	Weighted?	$C^{(2)}$
Authors as Nodes	PDS Nodes	No	0.981
	PDS Nodes	Yes	0.988
	Instrument Hosts	No	0.929
	Instrument Hosts	Yes	1.092
	Data Sets	No	0.936
	Data Sets	Yes	1.534
Events as Nodes	PDS Nodes	No	0.62
	PDS Nodes	Yes	1.927
	Instrument Hosts	No	0.715
	Instrument Hosts	Yes	1.675
	Data Sets	No	0.937
	Data Sets	Yes	1.455





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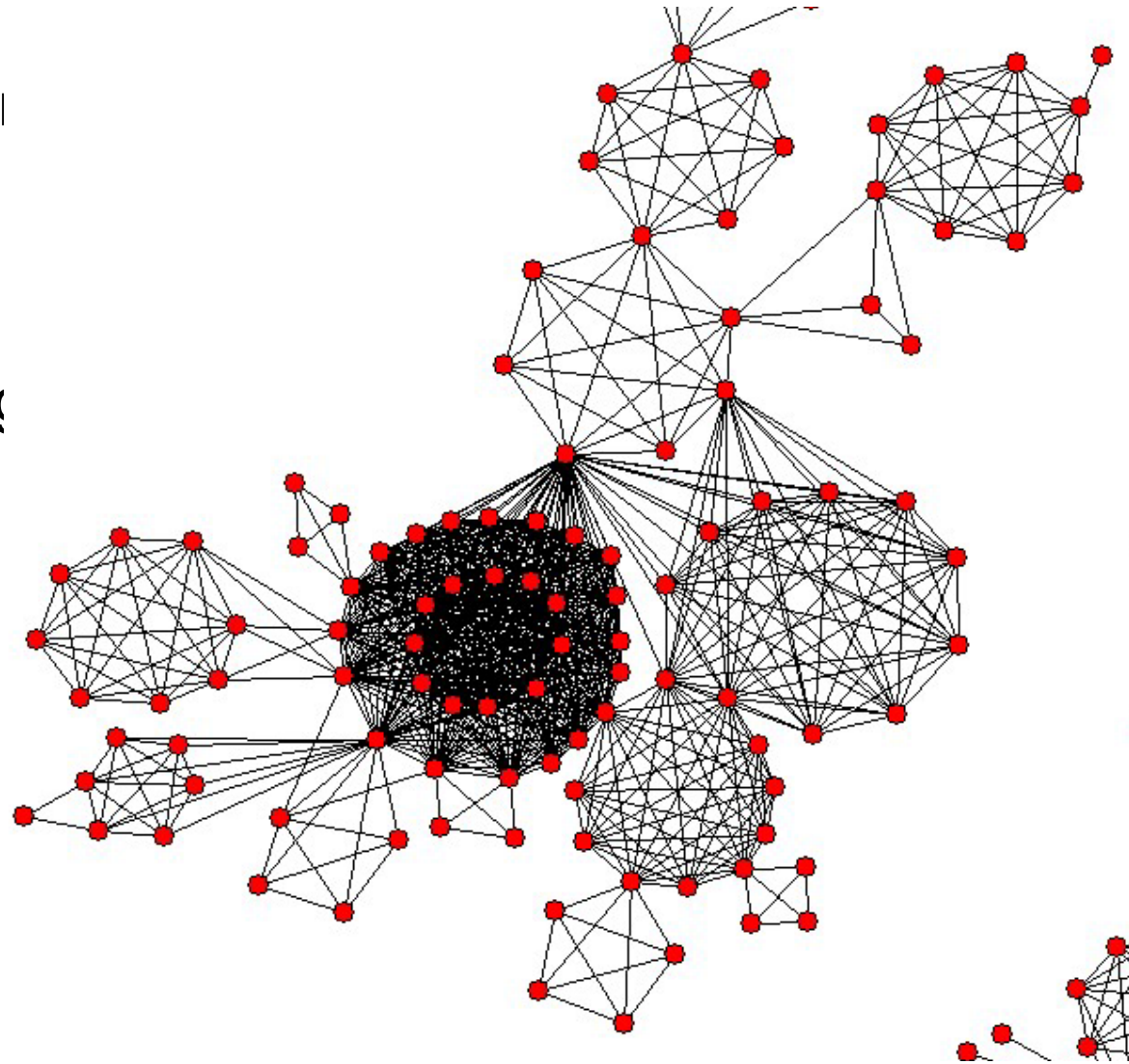
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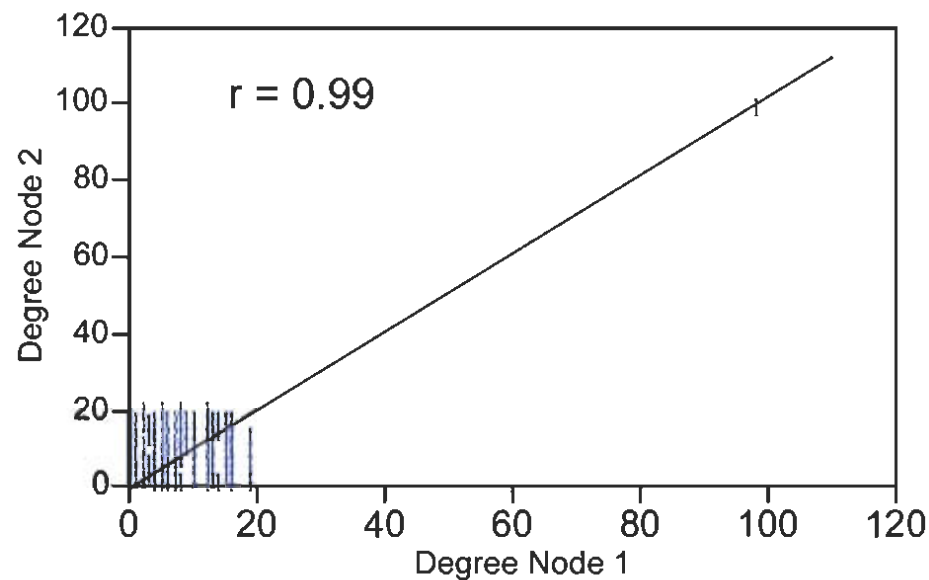
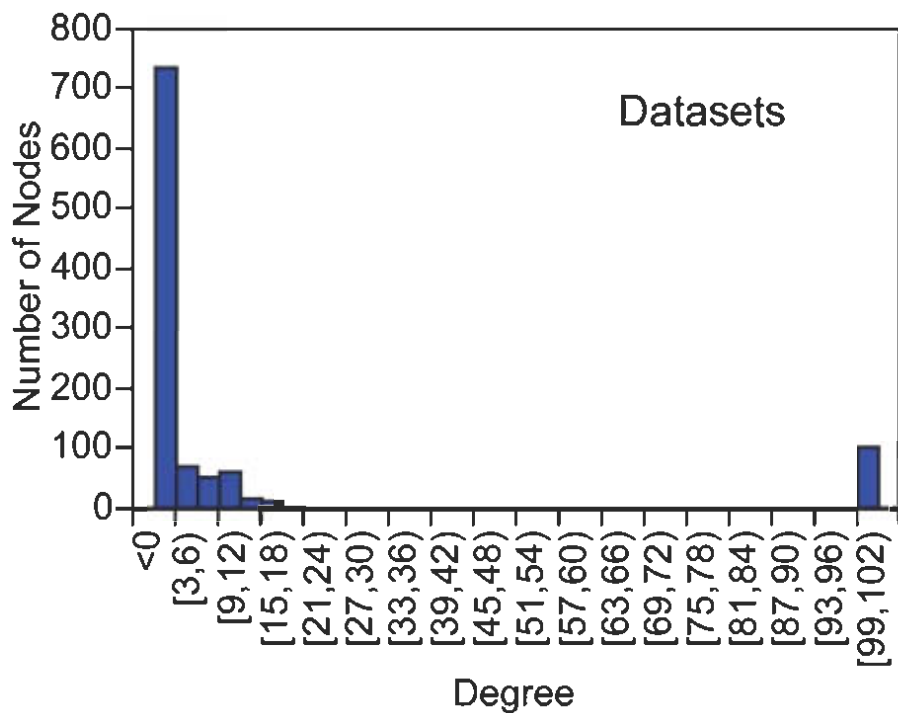
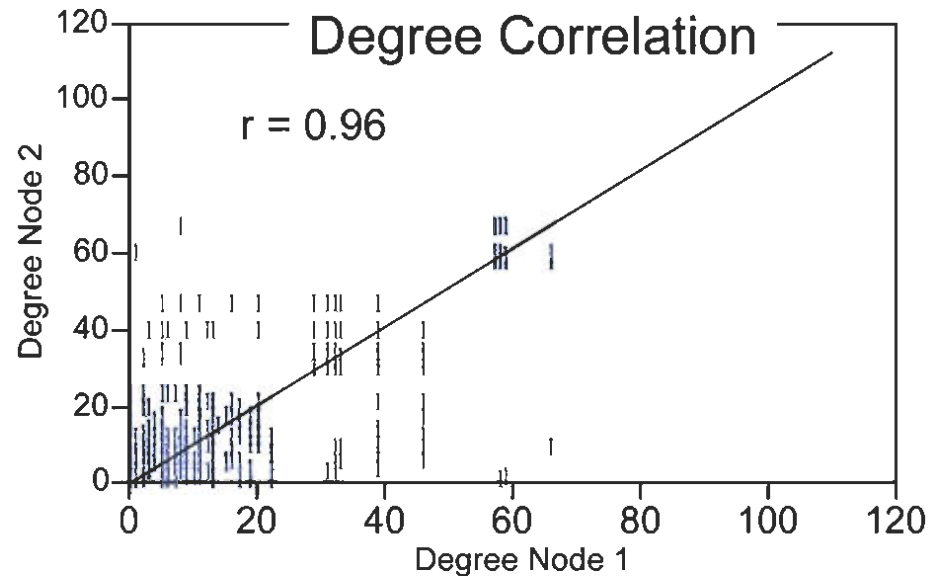
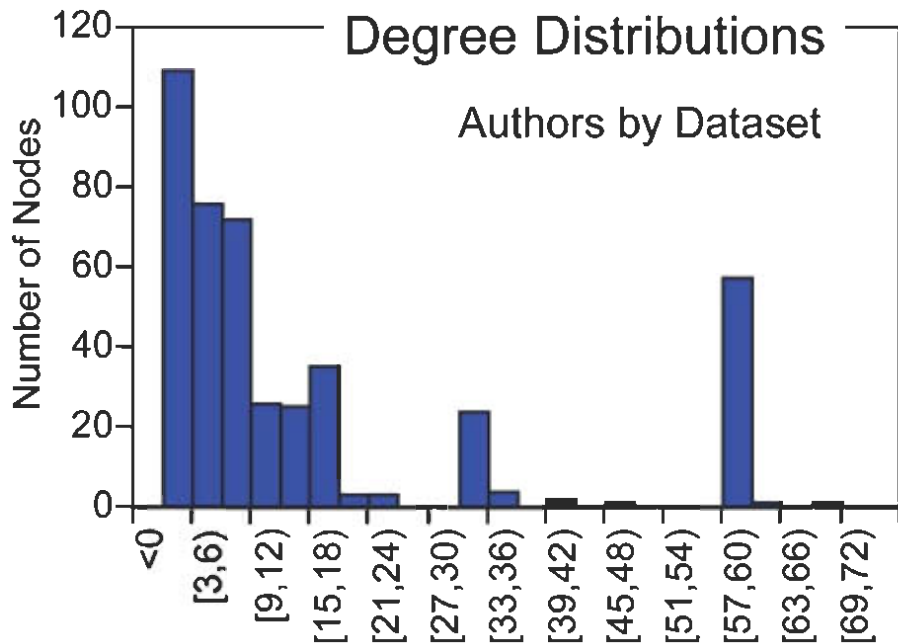
# Completely Connected Areas

==> High clustering coefficients

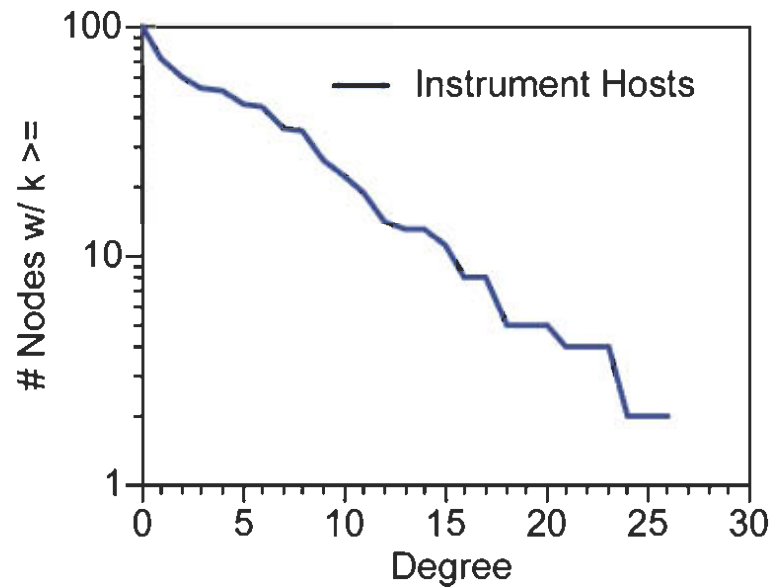
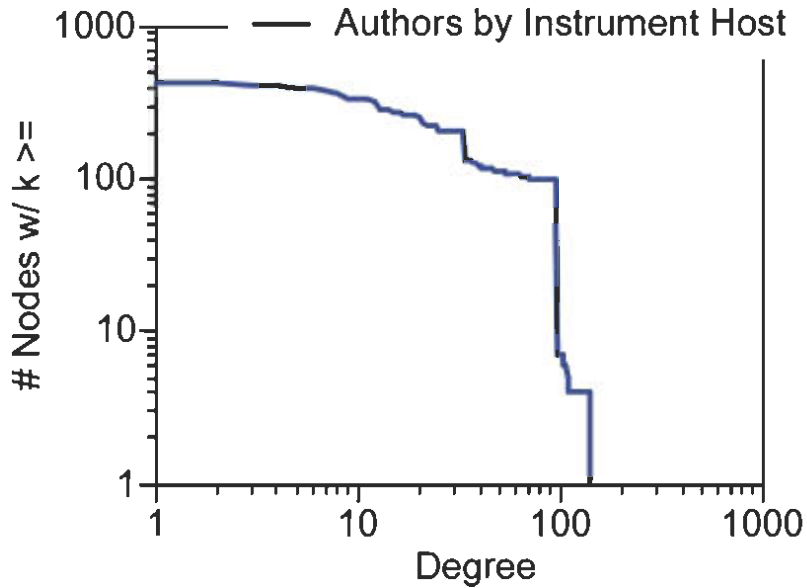
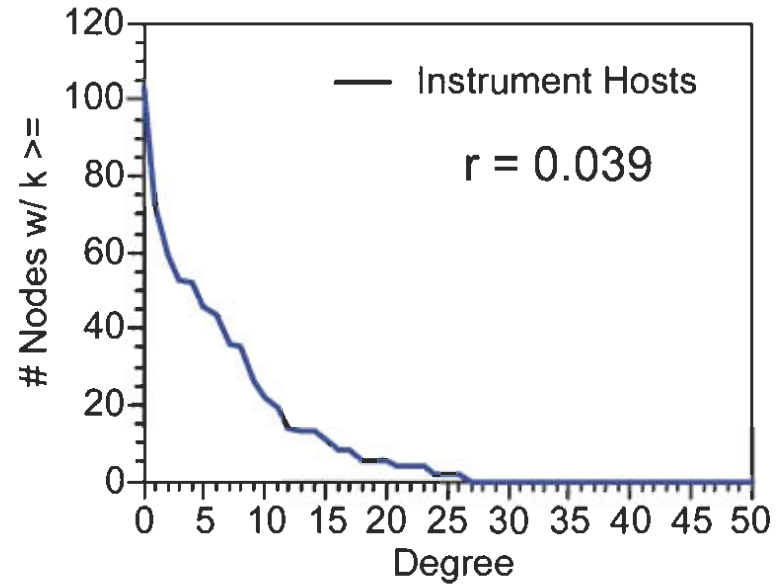
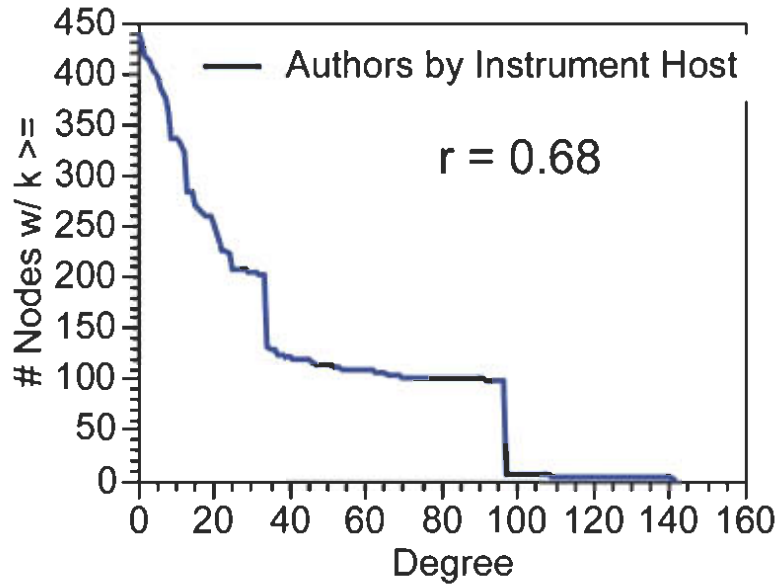
and

High Pearson degree correlation



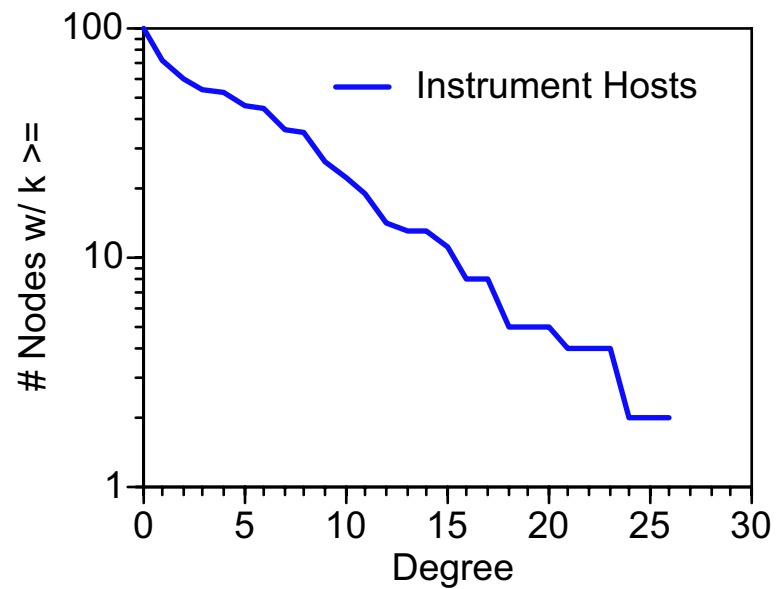
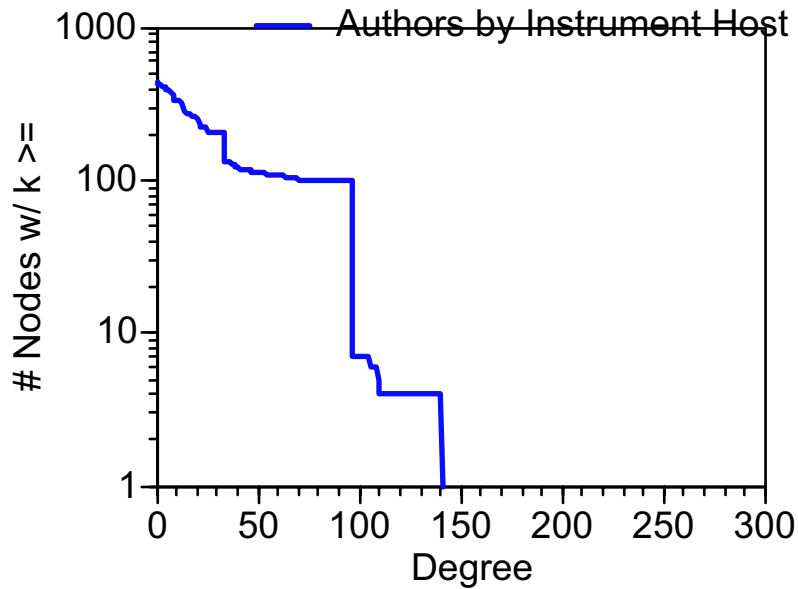
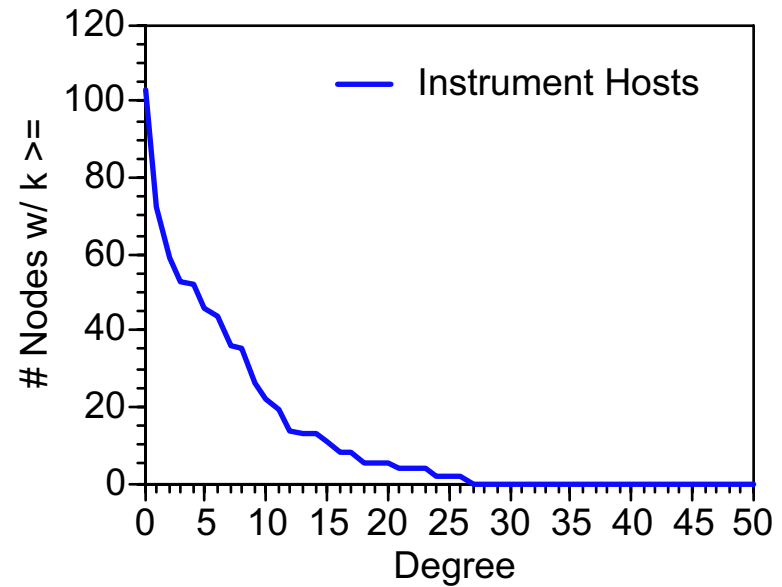
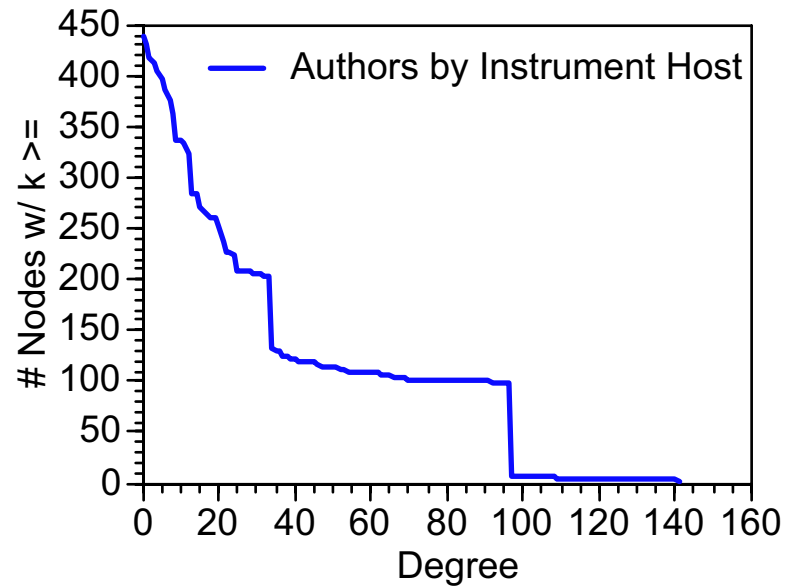


# Degree Distributions



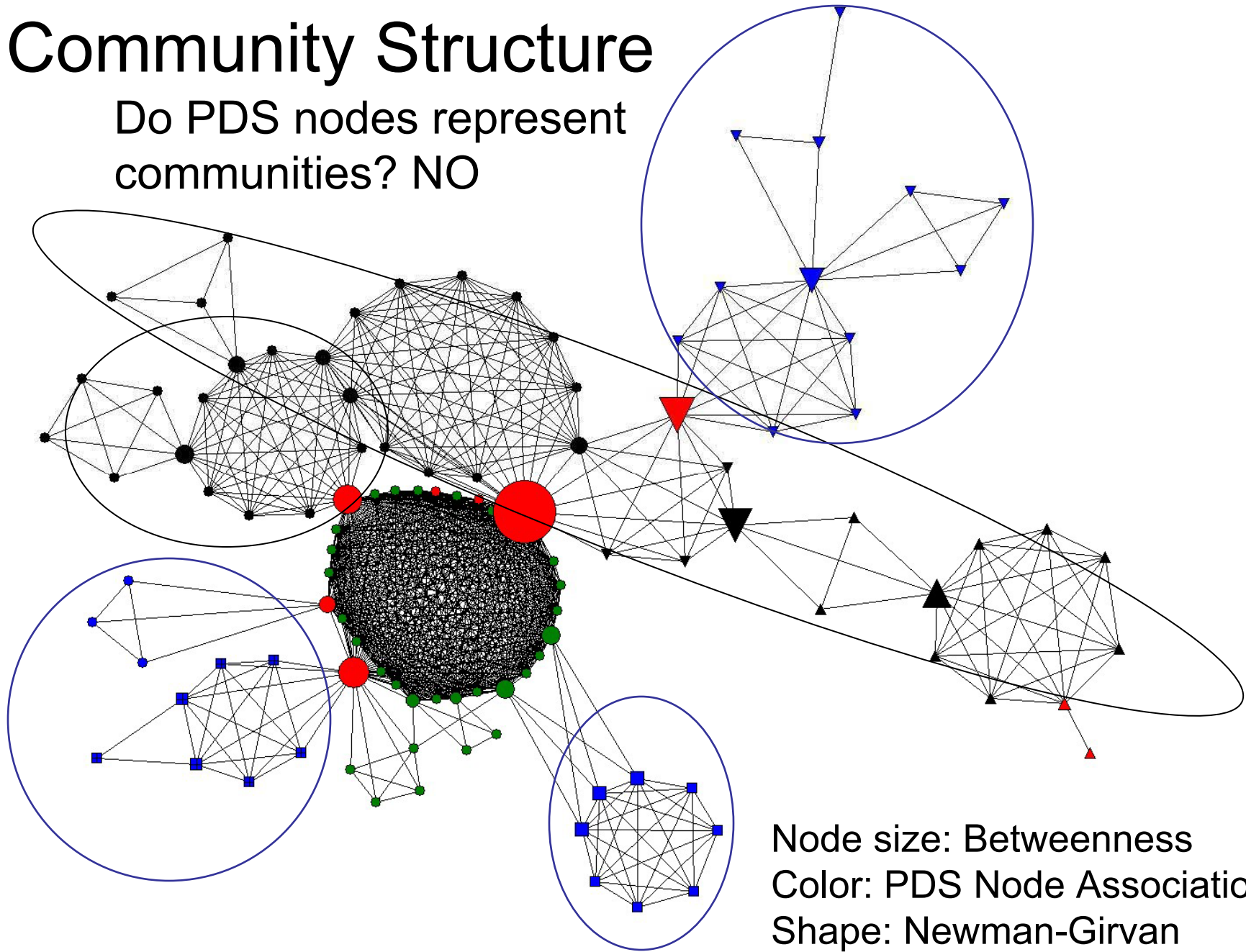


# Cumulative Degree Distributions



# Community Structure

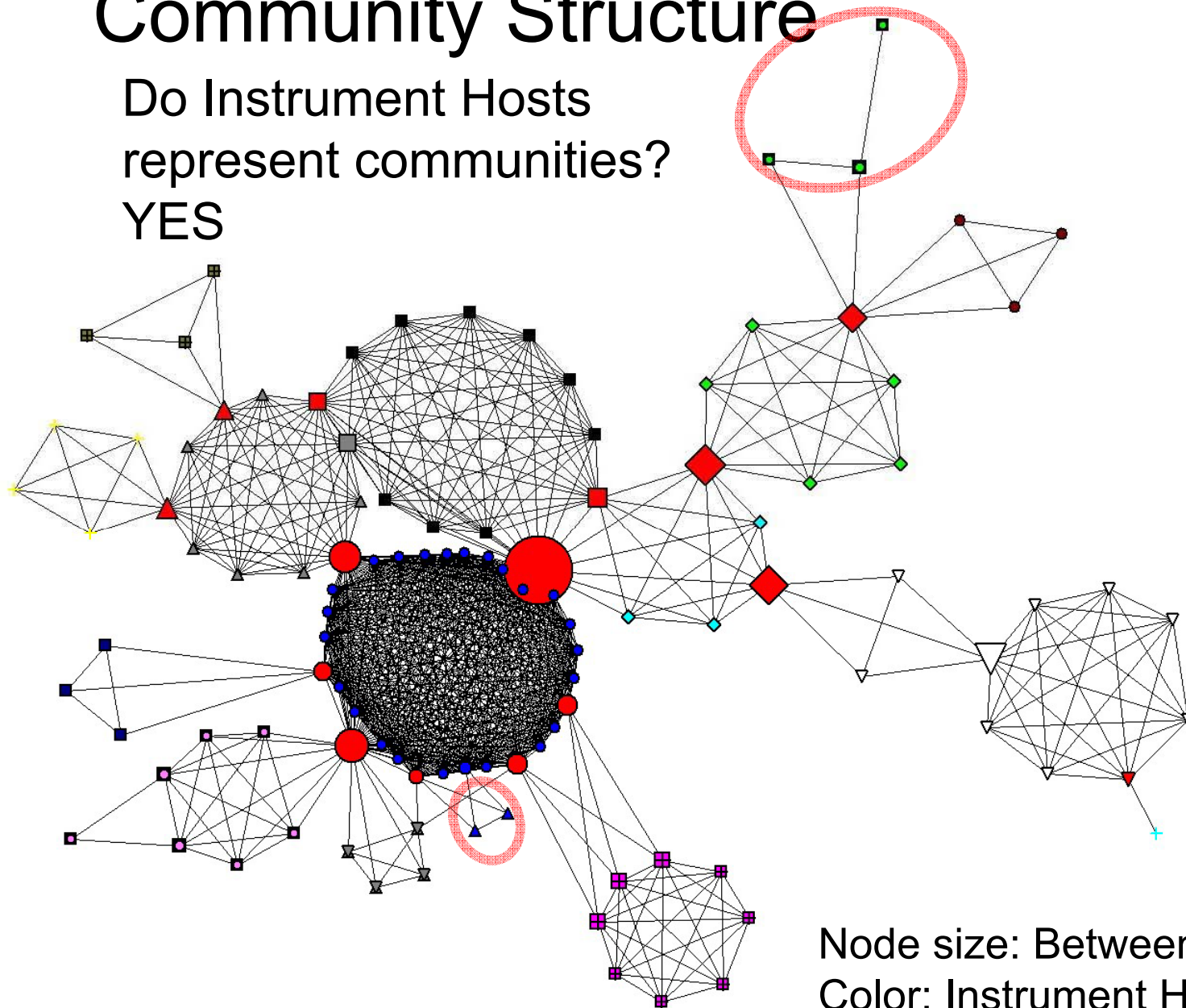
Do PDS nodes represent communities? NO



# Community Structure

Do Instrument Hosts  
represent communities?

YES

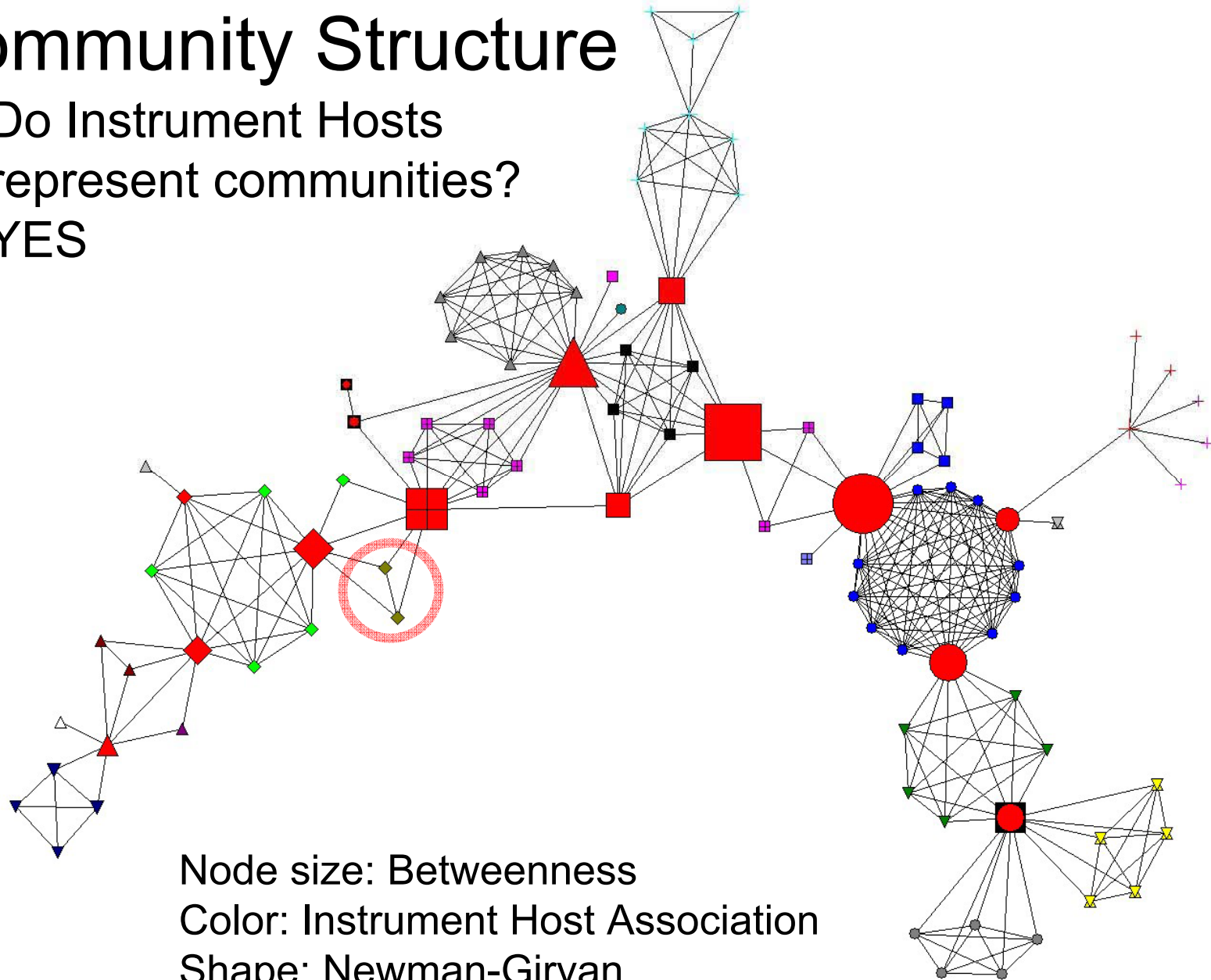


Node size: Betweenness  
Color: Instrument Host Association  
Shape: Newman-Girvan

# Community Structure

Do Instrument Hosts  
represent communities?

YES



# Centrality and Network Representation

Type of Node	Network	Weighted?	Centrality Measure			
			Degree	Closeness	Betweenness	Eigenvector
<b>Authors as Nodes</b>	PDS Nodes	No	0.339	0.364	0.136	0.060
	PDS Nodes	Yes	0.172	0.364	0.136	0.062
	Instrument Hosts	No	0.237	Unconnected	0.147	0.145
	Instrument Hosts	Yes	0.043	Unconnected	0.147	0.276
	Data Sets	No	0.120	Unconnected	0.026	0.354
	Data Sets	Yes	0.026	Unconnected	0.026	0.014
<b>Events as Nodes</b>	PDS Nodes	No	0.571	0.578	0.605	0.563
	PDS Nodes	Yes	0.314	0.578	0.605	2.152
	Instrument Hosts	No	0.205	Unconnected	0.074	0.385
	Instrument Hosts	Yes	0.046	Unconnected	0.074	1.555
	Data Sets	No	0.084	Unconnected	0.004	0.132
	Data Sets	Yes	0.006	Unconnected	0.004	0.512

- As before, weighting can affect the results.
  - Higher centrality without weighting.
  - Some measures are affected, and others are not.
- Some measures are not valid for some networks.
- Centrality tends to be highest for the network of PDS Nodes (with authors as edges).

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  - Small Bodies is a much better center without weighting.
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# What's Your Szego Number?

Overall Best Centers						
PDS Nodes	Instrument Hosts	Data Sets	Authors			
			By PDS Nodes	By Instrument Hosts	By Data Sets	Overall
Small_Bodies	Voyager 2	MER1-M-MI-5-	T. Z. Martin	C. Neese	K. Szego	K. Szego
Planetary_Atmospheres	Ground-Based Telescopes	MOSAIC-OPS-V1.0	R. Mehlman	L. S. Elson	J. T. Gosling	J. T. Gosling
	Public Literature		J. R. Spencer	C. H. Acton	R. F. Beebe	
				B.V.Semenov		

- Top two “best centers” determined for each of the 12 networks by each of the 4 metrics
- The most commonly appearing nodes selected as the “best centers” for each network (above)
  - Not surprisingly, two catch-alls are among the best centers (2<sup>nd</sup> and third) of instrument hosts
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# Future Work

- Look at the sub-network within one of the PDS nodes to determine centrality and community structure within a subfield.
- Consider PDS Data Access Statistics.
  - Only collected by NASA for access at the PDS node level by hostname (rather than individual). PDS staff may be able to collect data for data set and instrument host access, but it is not being done now.
  - We have collected the available usage statistics for some nodes (Small Bodies, Engineering, and PPI) and are in contact with all of the node managers.
- Create bipartite networks of authors and subject areas from a major planetary science conference.
  - American Astronomical Society's Division for Planetary Sciences (DPS) Annual Meeting has been suggested by R.P. Binzel (one of the authors here at MIT).
  - Compare results to those presented here. Are the community structure and centrality around instrument host and subject area specific to the PDS or more generalizable?
- Gather data on funding and political support for missions and for PDS nodes to determine the dynamics between scientists and policymakers and how that relates to the results presented here. (Possible component of Mark's research)